# The Normalization of the Resolution Function for Inelastic Neutron Scattering and its Application* 

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#### Abstract

By extending the considerations of Maier-Leibnitz, the normalization of the resolution function is found to be the product of the volumes in reciprocal space of the incident and scattered beams. Each volume is defined by an integration in reciprocal space over the probability of finding a particular $\mathbf{k}$, where $\mathbf{k}$ is the wave vector of the neutron. The resolution can be understood as a convolution of these two volumes. For three-axis spectrometers explicit expressions for these volumes are given. The knowledge of the normalization is necessary for numerical unfolding of experimental data. For two cases, which often occur in inelastic neutron scattering, it is possible to directly correct the experimental data without resorting to numerical unfolding. After applying these corrections the data represent the scattering law folded with a resolution function normalized to unity, i.e. the integral over the corrected data is the integral over the scattering law. It is shown that in this case, the unfolding of the corrected data turns out to be a one-dimensional problem.


## 1. Introduction

The resolution function of an instrument for inelastic neutron scattering depends on the wave vectors $\mathbf{k}_{i}$ and $\mathbf{k}_{f}$ of the incoming and scattered neutrons respectively. In the Gaussian approximation $\ddagger$ the surfaces of equal probability are four-dimensional ellipsoids. For a nominal setting $\mathbf{Q}_{0}$ and $\omega_{0}$ of the instrument the resolution function $R$ can be written

$$
\begin{align*}
& R\left(\mathbf{Q}-\mathbf{Q}_{0}, \omega-\omega_{0}\right)=R_{0}\left(\mathbf{Q}_{0}, \omega_{0}\right) \\
& \quad \times \exp \left\{-\frac{1}{2} \sum_{k=1}^{4} \sum_{l=1}^{4} M_{k t}\left(\mathbf{Q}_{0}, \omega_{0}\right) X_{k} X_{l}\right\} . \tag{1}
\end{align*}
$$

Here $\hbar \mathbf{Q}=\hbar \mathbf{k}_{i}-\hbar \mathbf{k}_{f}$ is the momentum transfer and $\hbar(1)$ $=\begin{gathered}\hbar^{2} \\ 2 m\end{gathered}\left(k_{i}^{2}-k_{f}^{2}\right)$ the energy transfer. $\hbar \mathbf{Q}_{0}$ and $\hbar \omega_{0}$ represent the most probable momentum and energy transfers and are defined in terms of the most probable values $\mathbf{k}_{I}$ and $\mathbf{k}_{F}$ of $\mathbf{k}_{i}$ and $\mathbf{k}_{f}$. The matrix $\mathbf{M}_{k l}$ describes the ellipsoids and the four coordinates $X_{k}$ represent the three com-

[^0]ponents of $\left(\mathbf{Q}-\mathbf{Q}_{0}\right)$ and $\omega-\omega_{0}$. The semi-major axes of these ellipsoids have been derived for a three-axis spectrometer by Stedman \& Nilsson (1966), Cooper \& Nathans (1967), Stedman (1968), Nielsen \& Bjerrum Møller (1969) and Bjerrum Møller \& Nielsen (1970) and for a time-of-flight instrument by Komura \& Cooper (1970). But the efficiency factor $R_{0}\left(\mathbf{Q}_{0}, \omega_{0}\right)$ and the normalization (by which we mean the integral over $R$ ) have not yet been discussed in the literature, although they play an important role in resolution corrections. For example, failure to take the normalization into consideration in analyzing experimental data can cause errors in the determination of both peak positions and line shapes. Integrated intensities, which are of increasing interest in inelastic neutron scattering measurements also depend crucially on the normalization.

In this paper a general and surprisingly simple expression for the normalization of the resolution is derived by extending Maier-Leibnitz's (1966) treatment of intensity and resolution in neutron scattering. Tucciarone, Lau,Corliss, Delapalme \& Hastings (1971) have also considered the question of normalizing resolution calculations. Their results are in agreement with those reported here after taking into account an improper normalization§ in Cooper \& Nathans' original paper. Neverthcless, in practice this error is

[^1]relatively minor and a computer program based on these considerations has been used to successfully analyze experimental data in the Brookhaven group for several years (Samuelsen, Hutchings \& Shirane, 1970; Lau, Corliss, Delapalme, Hastings, Nathans \& Tucciarone, 1970; Als-Nielsen, Axe \& Shirane, 1971; Hutchings, Shirane, Birgeneau \& Holt, 1971). This program and another one written by Hutchings and Samuelsen (1970), which also uses the erroneous expression of Cooper \& Nathans (Samuelsen, 1972), were found in good agreement with calibration measurements of phonons in copper (Shirane, 1971). Werner \& Pynn's (1971) calculations of corrections for resolution effects follow a different approach and do not give an explicit normalization. Numerical integration was checked to agree with the normalization given here (Werner, 1971).

The authors cited in the last paragraph have only considered resolution effects in three-axis spectrometers. Dietrich (1971) has treated the analogous problem for a time-of-flight spectrometer and has derived not only the parameters of the resolution function but the normalization as well.

The paper is organized in the following way. In Section 2 the counting rate at the detector is calculated following the 'phase space' considerations of MaierLeibnitz. This approach is the key to a transparent interpretation of the normalization of the resolution function, which is derived in Section 3. An explicit expression of the normalization for a three-axis spectrometer is given in Section 4. In Section 5 we discuss the corrections to experimental data obtained with a three-axis spectrometer and in Section 6 we show that by using the Gaussian approximation in calculating the resolution function the unfolding of the data often becomes a one-dimensional problem.

## 2. The neutron counting rate at the detector

From equations (5) and (7) of Maier-Leibnitz we get the infinitesimal current incident on the sample

$$
\begin{equation*}
\mathrm{d} J_{i}=A^{\prime}\left(k_{i}\right) \cdot \mathrm{d} F \cdot k_{i} \cdot p_{i}\left(\mathbf{k}_{i}\right) \cdot \mathrm{d} k_{i x} \cdot \mathrm{~d} k_{i y} \cdot \mathrm{~d} k_{i z} \tag{2}
\end{equation*}
$$

where $\mathrm{d} F$ is an element of the geometric beam cross section. For a sample small compared to the beam area, the current can be assumed constant and $\mathrm{d} F$ can be integrated to $F$, which is the area of the sample. $p_{i}\left(\mathbf{k}_{i}\right)$ is the probability of finding a neutron with wave vector $\mathbf{k}_{i}$ in the beam produced by any monochromator and incident on the sample. $p_{i}\left(\mathbf{k}_{i}\right)$ replaces the quantity $f_{o}$ used by Maier-Leibnitz.

Under the assumption of a Maxwell distribution in the reactor we write

$$
A^{\prime}\left(k_{i}\right)=\frac{\varphi}{2 \pi k_{T}^{4}} \exp \left(-\begin{array}{c}
k_{i}^{2} \\
k_{T}^{2}
\end{array}\right),
$$

where $\varphi$ is the thermal flux in the reactor,$k_{T}=$
$\left[\frac{2 m}{\hbar^{2}} k_{B} T\right]^{1 / 2}, k_{B}$ is the Boltzmann constant and $T$ the effective temperature of the moderator. Fig. 1 shows schematically the scattering in real and in reciprocal space.

The scattering cross section of the sample is

$$
\begin{equation*}
\frac{\mathrm{d}^{2} \Sigma}{\mathrm{~d} \Omega \mathrm{~d} \omega}=N \frac{k_{f}}{k_{i}} S(\mathbf{Q}, \omega) \tag{3}
\end{equation*}
$$

$\Sigma$ is the macroscopic scattering cross section i.e. the scattering probability per cm and per unit area of the sample and $N$ is the density of unit cells in the sample By equation (3) we define a scattering law $S(\mathbf{Q}, \omega)$, which contains the scattering lengths of all nuclei in the sample.

In order to derive the infinitesimal current at the detector we use elements of the scattered beam in units,


Fig. 1. Schematic drawing of an inelastic neutron scattering experiment: (a) in real space, $(b)$ in reciprocal space.


Fig. 2. Distribution of monochromatic neutrons in the scattering plane produced by a single crystal monochromator, where $0_{M}$ is the Bragg angle, $2 \pi \tau_{M}$ the rec. lattice vector of the reflecting planes, $\alpha_{0}$ and $\alpha_{1}$ are the horizontal collimations of the incoming and outgoing beams and $\eta_{M}$ is the horizontal mosaic width of the crystal.
which are compatible with the cross section given in equation (3). These elements are

$$
\mathrm{d} \Omega_{f}=\frac{\mathrm{d} k_{f x \cdot} \cdot \mathrm{~d} k_{f y}}{k_{f}^{2}}, \quad \mathrm{~d} \omega_{f}=\frac{\hbar}{m} k_{f} \mathrm{~d} k_{f_{z}} .
$$

Let $p_{f}\left(\mathbf{k}_{f}\right)$ be the probability that a neutron with wave vector $k_{f}$ passes the analyzer and is counted at the detector. Then for a sample of thickness $t$ the infinitesimal current detected is

$$
\begin{align*}
\mathrm{d} J^{\prime} & =\mathrm{d} J_{i} \cdot \frac{\mathrm{~d}^{2} \Sigma}{\mathrm{~d} \Omega \mathrm{~d} \omega} \cdot t \cdot p_{f}\left(\mathbf{k}_{f}\right) \cdot \mathrm{d} \Omega_{f} \cdot \mathrm{~d} \omega_{f} \\
& =A \cdot S(\mathbf{Q}, \omega) \cdot p_{i}\left(\mathbf{k}_{i}\right) \cdot p_{f}\left(\mathbf{k}_{f}\right) \cdot \mathrm{d} k_{i x} \cdot \mathrm{~d} k_{i y} \cdot \mathrm{~d} k_{i z} \\
& \times \mathrm{d} j_{f x} \cdot \mathrm{~d} k_{f y} \cdot \mathrm{~d} k_{f z} \tag{4}
\end{align*}
$$

where $A$ includes among other factors $A^{\prime}\left(k_{i}\right), t$ and the sample area $F$. By substituting

$$
\begin{gather*}
\mathrm{d} V_{i}=\mathrm{d} k_{i x} \cdot \mathrm{~d} k_{i y} \cdot \mathrm{~d} k_{i z} \\
d V_{f}=\mathrm{d} k_{f x} \cdot \mathrm{~d} k_{f y} \cdot \mathrm{~d} k_{f z} \tag{5}
\end{gather*}
$$

in equation (4) one sees that the infinitesimal current at the detector is proportional to the volume elements in $k$ space for the incoming and outgoing beam, as Maier-Leibnitz has shown.

Equation (4) is not as simple as it appears since the scattering law depends on $\mathbf{Q}$ and $\omega$, while the independent variables in equation (4) are $\mathbf{k}_{i}$ and $\mathbf{k}_{f}$. Therefore, we collect all combinations of $\mathbf{k}_{\boldsymbol{i}}$ with $\mathbf{k}_{f}$, which contribute to a given $\mathbf{Q}$ and $\omega$, and write $\mathrm{d} J$ proportional to elements in $\mathbf{Q}, \omega$ space;

$$
\begin{align*}
\mathrm{d} J & =A \cdot S(\mathbf{Q}, \omega) \iint p_{i}\left(\mathbf{k}_{i}\right) \cdot p_{f}\left(\mathbf{k}_{f}\right) \cdot \delta\left[\mathbf{Q}-\left(\mathbf{k}_{i}-\mathbf{k}_{f}\right)\right] \\
& \times \delta\left[\omega-\frac{\hbar}{2 m}\left(k_{i}^{2}-k_{f}^{2}\right)\right] \cdot \mathrm{d} V_{i} \cdot \mathrm{~d} V_{f} \cdot \mathrm{~d}^{3} Q \cdot \mathrm{~d} \omega \tag{6}
\end{align*}
$$

Both $\mathrm{d} J$ and $\mathrm{d} J^{\prime}$ have the dimension neutrons per sec. Although $\int \mathrm{d} J^{\prime}=\int \mathrm{d} J, \mathrm{~d} J^{\prime}$ is defined in elements of $\mathbf{k}_{i} . \mathbf{k}_{f}$ whereas $\mathrm{d} J$ is written for clements of $\mathbf{Q}$ and $\omega$. The $\delta$ functions in (6) represent conservation of momentum and energy. We can rewrite equation (6) in a more condensed form

$$
\begin{equation*}
d J=A \cdot S(\mathbf{Q}, \omega) \cdot R^{\prime}(\mathbf{Q}, \omega) \cdot \mathrm{d}^{3} Q \cdot \mathrm{~d} \omega \tag{7}
\end{equation*}
$$

where the scattering law $S(\mathbf{Q}, \omega)$ describes the physical characteristics of the sample and $R^{\prime}(\mathbf{Q}, \omega)$ the transmission or resolution function of the instrument. A particular setting of the instrument can be described by the most probable values $\mathbf{k}_{I}$ and $\mathbf{k}_{F}$ or preferably by $\mathbf{Q}_{0}$ and $\omega_{0}$. The resolution function $R^{\prime}(\mathbf{Q}, \omega)$ can be redefined in terms of $\mathbf{Q}_{0}$ and $\omega_{0}$

$$
\begin{align*}
R^{\prime}(\mathbf{Q}, \omega) & \equiv R\left(\mathbf{Q}-\mathbf{Q}_{0}, \omega-\omega_{0}\right) \\
& =\iint \delta\left[\left(\mathbf{Q}-\mathbf{Q}_{0}\right)-\left\{\mathbf{k}_{i}-\mathbf{k}_{f}-\left(\mathbf{k}_{I}-\mathbf{k}_{F}\right)\right\}\right] \\
& \times \delta\left[\left(\omega-\omega_{0}\right)-\frac{\hbar}{2 m}\left\{\mathbf{k}_{i}^{2}-\mathbf{k}_{f}^{2}-\left(\mathbf{k}_{I}^{2}-\mathbf{k}_{F}^{2}\right)\right\}\right] \\
& \times p_{i}\left(\mathbf{k}_{i}\right) \cdot p_{f}\left(\mathbf{k}_{f}\right) \cdot \mathrm{d} V_{i} \cdot \mathrm{~d} V_{f} \tag{8}
\end{align*}
$$

Thus the counting rate at the detector for a position $\mathbf{Q}_{0}, \omega_{0}$ is given by the expression

$$
\begin{equation*}
J\left(\mathbf{Q}_{0}, \omega_{0}\right)=A \int S(\mathbf{Q}, \omega) . R\left(\mathbf{Q}-\mathbf{Q}_{0}, \omega-\omega_{0}\right) \mathrm{d}^{3} Q \mathrm{~d} \omega \tag{9}
\end{equation*}
$$

Note that our definition of the resolution function $R$ differs from that of Cooper \& Nathans by the factor $k_{f} / k_{i}$, which we have included. This definition has two advantages; first, $R$ is now easier to understand as explained in the next section and second, the integration in equation (9) is performed over the most suitable variables, $\mathbf{Q}$ and $\omega$. As for scattering of thermal neutrons the scattering law $S(\mathbf{Q}, \omega)$ depends only on the momentum transfer $\hbar \mathbf{Q}$ and the energy transfer $\hbar \omega$ and does not depend on the chosen pair of $k_{i}$ and $k_{f}$, an integration over these latter quantities as performed by Collins (1963) is unnecessarily complicated and obscures the clear separation of sample from instrumental properties as shown in equation (9).

## 3. The resolution function

To visualize the resolution function, it is convenient to show that $R$ as a function of $\mathbf{Q}$ is a convolution of the probability $p_{i}$ with the probability $p_{f}$. Therefore we integrate $R^{\prime}(\mathbf{Q}, \omega)$ [see equations (6) and (7)] over $\omega$ and obtain

$$
\begin{align*}
& R^{\prime \prime}(\mathbf{Q})=\int R^{\prime}(\mathbf{Q}, \omega) \mathrm{d} \omega \\
= & \int_{V_{I}} \int_{V_{F}} \delta\left[\mathbf{Q}-\left(\mathbf{k}_{i}-\mathbf{k}_{f}\right)\right] \cdot p_{i}\left(\mathbf{k}_{i}\right) \cdot p_{f}\left(\mathbf{k}_{f}\right) \cdot \mathrm{d} V_{i} \cdot \mathrm{~d} V_{f} \tag{10}
\end{align*}
$$

The probabilities $p_{i}$ and $p_{f}$ describe only neutrons traveling in the direction of the beam. When $k_{z}$ is negative ( $z$ being the direction of the beam) $p_{i}$ and $p_{f}$ are zero. This enables us to integrate from $-\infty$ to $+\infty$. The integration of equation (10) is equivalent to a projection of $R^{\prime}$ onto the $\omega=0$ plane (see Section 6).

After integration over $V_{F}$ we have

$$
\begin{equation*}
R^{\prime \prime}(\mathbf{Q})=\int_{V_{I}} p_{i}\left(\mathbf{k}_{i}\right) p_{f}\left(\mathbf{k}_{i}-\mathbf{Q}\right) \mathrm{d} V_{i} \tag{11}
\end{equation*}
$$

which shows that $R^{\prime \prime}(\mathbf{Q})$ is a convolution of the distri-
bution of $\mathbf{k}_{i}$ within $V_{I}$ with the distribution of $\mathbf{k}_{\boldsymbol{f}}$ within $V_{F}$.

Alternatively, if we integrate $R^{\prime}(\mathbf{Q}, \omega)$ over $\mathbf{Q}$ we get

$$
\begin{align*}
R^{\prime \prime \prime}(\omega) & =\int R^{\prime}(\mathbf{Q}, \omega) \mathrm{d}^{3} Q \\
& =\int_{V_{I}} \int_{V_{F}} \delta\left[\omega-\frac{\hbar}{2 m}\left(k_{i}^{2}-k_{f}^{2}\right)\right] \\
& \times p_{i}\left(\mathbf{k}_{i}\right) \cdot p_{f}\left(\mathbf{k}_{f}\right) \cdot \mathrm{d} V_{i} \cdot \mathrm{~d} V_{f} \tag{12}
\end{align*}
$$

The problem here is that the $\delta$-function is a condition on the energy, which means that it selects particular combinations of $k_{i}^{2}$ and $k_{f}^{2}$, whereas the integration is performed in $k$-space, that is with equal steps in $k$.

First we define a new quantity $\operatorname{prob}\left(k_{z}\right)$ obtained by integrating over $k_{x}$ and $k_{y} . k_{z}$ is parallel to the most probable $\mathbf{k}\left(\mathbf{k}_{I}\right.$ or $\mathbf{k}_{F}$ )

$$
\begin{equation*}
\operatorname{prob}\left(k_{z}\right)=\int p(\mathbf{k}) \mathrm{d} k_{x} \mathrm{~d} k_{y} \tag{13}
\end{equation*}
$$

After integration of equation (12) over the $x$ and $y$ components and $\mathrm{d} k_{f z}$ we get
$R^{\prime \prime \prime}(\omega)=\int \operatorname{prob}_{i}\left(k_{i z}\right) \cdot \operatorname{prob}_{f}\left[\sqrt{ }\left(k_{i z}^{2}-\frac{2 m}{\hbar} \cdot \omega\right)\right] \mathrm{d} k_{i z}$.

While previously $p$ was defined to be non-zero only when $k_{z}$ is positive, now we must define $\operatorname{prob}\left(1 / k_{z}^{2}\right)$ to be non-zero only for positive $k^{2}$. Again we can integrate from $-\infty$ to $+\infty$. It is easily seen that equation (14) turns out to be a convolution in energy

$$
\begin{align*}
R^{\prime \prime \prime}(\omega) & =\int \operatorname{prob}_{i}\left(\gamma k_{i z}^{2}\right) \cdot \operatorname{prob}_{f}\left[\sqrt{ }\left(k_{i z}^{2}-\begin{array}{c}
2 m \\
\hbar
\end{array} \omega\right)\right] \\
& \times{ }_{2 k_{i z}}^{1} \mathrm{~d}\left(k_{i z}\right) \tag{15}
\end{align*}
$$

But equation (15) is not just a convolution of the probabilities of finding a particular $k_{i z}$ or a $k_{f z}$, but rather the convolution of the probabilities weighted by $1 / k_{i z}$. Another interpretation is that (15) is a convolution of the two probabilities with respect to energy, but the step width is weighted to get equal steps in $k$ space rather than in $k^{2}$ or energy space.

The properties of the resolution function discussed so far in this section are intended to help visualize the contributions from various components. We will make use of this convolution consideration in Section 5.

Let us now turn to the very important question of normalization. This requires that we integrate equation (10) over $\mathbf{Q}$

$$
\begin{array}{r}
\iint R^{\prime}(\mathbf{Q}, \omega) \mathrm{d}^{3} Q \mathrm{~d} \omega=\iint R\left(\mathbf{Q}-\mathbf{Q}_{0}, \omega-\omega_{0}\right) \mathrm{d}^{3} Q \mathrm{~d} \omega \\
\quad=\int_{V_{I}} \int_{V_{F}} p_{i}\left(\mathbf{k}_{i}\right) \cdot p_{f}\left(\mathbf{k}_{f}\right) \mathrm{d} V_{i} \mathrm{~d} V_{f}=V_{I} \cdot V_{F} \tag{16}
\end{array}
$$

with

$$
\begin{equation*}
\int_{V_{I}} p_{i}\left(\mathbf{k}_{i}\right) \mathrm{d} V_{i}=V_{I} \tag{17}
\end{equation*}
$$

and

$$
\int_{V_{F}} p_{f}\left(\mathbf{k}_{f}\right) \mathrm{d} V_{f}=V_{F}
$$

The normalization is proportional to the product of the volumes [as defined by (17)] in reciprocal space: $V_{I}$ around $\mathbf{k}_{I}$ and $V_{F}$ around $\mathbf{k}_{F}$, and is independent of the scattering angle. Equation (16) contains no approximations and holds for three-axis spectrometers as well as for time-of-flight instruments.

## 4. The volumes $\boldsymbol{V}_{\boldsymbol{I}}$ and $\boldsymbol{V}_{\boldsymbol{F}}$ for single-crystal systems

From the derivation given in the Appendix, we find

$$
\begin{align*}
V_{I} & =P_{M}\left(k_{I}\right) \cdot k_{I}^{3} \cdot \cot \theta_{M} \cdot(2 \pi)^{3 / 2} \\
& \times \frac{\beta_{0} \beta_{1}}{V\left(4 \sin ^{2}\left(\theta_{M}\right) \eta_{M}^{\prime}{ }^{2}+\beta_{0}^{2}+\beta_{1}^{2}\right)} \\
& \times \frac{\eta_{M} \cdot \alpha_{0} \cdot \alpha_{1}}{V\left(\alpha_{0}^{2}+\alpha_{1}^{2}+4 \eta_{M}^{2}\right)}  \tag{18a}\\
& \times \frac{P_{A}\left(k_{F}\right) \cdot k_{F}^{3} \cdot \cot \theta_{A} \cdot(2 \pi)^{3 / 2}}{V\left(4 \sin ^{2}\left(\theta_{A}\right) \eta_{A}^{2}+\beta_{2}^{2}+\beta_{3}^{2}\right)} \\
& \times \frac{\beta_{A} \cdot \alpha_{2} \cdot \alpha_{3}}{V\left(\alpha_{2}^{2}+\alpha_{3}^{2}+4 \eta_{A}^{2}\right)} .
\end{align*}
$$



Fig. 3. Inelastic neutron scattering results of $\mathrm{Tb}_{2}\left(\mathrm{MoO}_{4}\right)_{3}$. The solid line represents a least-squares fit of a symmetric scattering law multiplied with the normalization of the resolution function.
$\alpha_{0}, \alpha_{1}, \alpha_{2}, \alpha_{3}$ and $\beta_{0}, \beta_{1}, \beta_{2}, \beta_{3}$ are the horizontal and vertical divergencies of the collimators, where $0,1,2,3$ refer to inpile, monochromator-to-sample, sample-toanalyzer and analyzer-to-detector respectively, $\eta_{M}, \eta_{M}^{\prime}$, and $\eta_{A}, \eta_{A}^{\prime}$ are the horizontal and vertical mosaic widths of the monochromating and analyzing crystals respectively and $\theta_{M}, \theta_{A}$ are the Bragg angles of monochromator and analyzer.

In many cases [ $\left.4 \sin ^{2}\left(\theta_{M}\right) \eta_{M}^{\prime 2}\right]$ can be neglected compared to ( $\beta_{0}^{2}+\beta_{1}^{2}$ ). Further, if $P_{M}\left(k_{I}\right)$ and $P_{A}\left(k_{F}\right)$ are assumed to be constant, one gets

$$
\begin{align*}
V_{I} & \propto k_{I}^{3} \cdot \cot \theta_{M}  \tag{19a}\\
V_{F} & \propto k_{F}^{3} \cdot \cot \theta_{A} . \tag{19b}
\end{align*}
$$

It is easy to understand relations (19a) and (19b) by referring to Fig. 2 which shows $V_{I}$ projected onto the scattering plane for $\alpha_{0}>\alpha_{1}$. The projection of $V_{1}$ is proportional to $\left(\eta_{M}, k_{I} \cdot \cos \theta_{M}\right) \cdot\left(\alpha_{1}, k_{I} / \sin \theta_{M}\right)$. For $\beta_{0}>\beta_{1}$ the vertical contribution (normal to the plane of the Fig. 2) gives a factor $k_{I} \cdot \beta_{1}$. It then follows that

$$
V_{I} \simeq k_{I}^{3} \cdot \frac{\cos \theta_{M}}{\sin \theta_{M}} \cdot \eta_{M} \cdot \alpha_{1} \cdot \beta_{1} .
$$

Besides $P_{M}\left(k_{I}\right)$ and $P_{A}\left(k_{F}\right), \eta_{M}, \eta_{M}^{\prime}, \eta_{A}$ and $\eta_{A}^{\prime}$ also depend on $k_{I}$ and $k_{F}$ respectively (Dorner, 1971) but the dependence on the wave vector can be measured and incorporated empirically into the expressions (18a) and ( $18 b$ ).

The above expressions for $V_{I}$ and $V_{F}$ contain no additional approximations beyond those made by Cooper \& Nathans, i.e. the widths of $p_{i}\left(\mathbf{k}_{i}\right)$ and $p_{f}\left(\mathbf{k}_{f}\right)$ are small compared to $\mathbf{k}_{I}$ and $\mathbf{k}_{F}$ respectively and


Fig.4. The solid line represents the true effect of the resolution for varying $\eta_{1}$ and the broken line gives the approximation discussed in Section 5.

Gaussian distributions representing collimators and mosaic widths.

## 5. Intensity corrections for resolution effects

In Section 3 the resolution function has been derived as a property of the measuring instrument. In other words, $R\left(\mathbf{Q}-\mathbf{Q}_{0}, \omega-\omega_{0}\right)$ describes the characteristics of the instrument at a nominal position $\mathbf{Q}_{0}, \omega_{0}$. In this section, we are going to combine the characteristics of the sample, given by $S(\mathbf{Q}, \omega)$, with the resolution function $R$. From equation (9) we see that the counting rate $J$ is a folding of $S$ and $R$. There is a complication in that the four-dimensional Gaussian $R$ has different widths and heights at different $\mathbf{Q}_{0}$ and $\omega_{0}$, as can be seen from equation (1). If we diagonalize the matrix $M_{k l}\left(\mathbf{Q}_{0}, \omega_{0}\right)$ to $M_{k l}^{\prime}\left(\mathbf{Q}_{0}, \omega_{0}\right)$, then we can find the widths $w_{k}\left(\mathbf{Q}_{0}, \omega_{0}\right)$ in the directions of the four principal axes by using the relation

$$
\begin{equation*}
\frac{1}{2} M_{k l}^{\prime}\left(\mathbf{Q}_{0}, \omega_{0}\right)=\frac{1}{w_{k}^{2}} \overline{\left(\mathbf{Q}_{0}, \omega_{0}\right)} \delta_{k l} \tag{20}
\end{equation*}
$$

We integrate equation (1) and use equations (16) and (20) to obtain

$$
\begin{align*}
\int R(\mathbf{Q} & \left.-\mathbf{Q}_{0}, \omega-\omega_{0}\right) \mathrm{d}^{3} Q \mathrm{~d} \omega \\
& =R_{0}\left(\mathbf{Q}_{0}, \omega_{0}\right) \pi^{2} \cdot \prod_{k=1}^{4} w_{k}\left(\mathbf{Q}_{0}, \omega_{0}\right)=V_{I} . V_{F} \tag{21}
\end{align*}
$$

The efficiency factor $R_{0}$ now can be expressed in terms of the normalization and the widths

$$
\begin{equation*}
R_{0}\left(\mathbf{Q}_{0}, \omega_{0}\right)=\frac{V_{1} \cdot V_{F}}{\pi^{2} \cdot \prod_{k=1}^{4} \cdot w_{k}\left(\mathbf{Q}_{0}, \omega_{0}\right)} \tag{22}
\end{equation*}
$$

As mentioned in the introduction, the $w_{k}$ for a threeaxis spectrometer have been calculated by several authors. $V_{I}$ and $V_{F}$ for such an instrument are given in Section 4 of this paper. Hence equation (22) fully determines $R_{0}\left(\mathbf{Q}_{0}, \omega_{0}\right)$.
$R_{0}$ is the quantity which we must have in order to numerically integrate equation (9), i.e. to fold our resolution function with the scattering law. There are however two situations which are commonly encountered in inelastic neutron scattering measurements in which it is not necessary to do the numerical integration when the product $V_{I} . V_{F}$ is known. Let us consider these in more detail.

## (a) $S(\mathbf{Q}, \omega)$ slowly varying in all four dimensions

By slowly varying we mean that $S$ varies linearly within the resolution volume. Then, since $R$ (within Gaussian approximation) is an even function, $S(\mathbf{Q}, \omega)$ can be replaced by $S\left(\mathbf{Q}_{0}, \omega_{0}\right)$ and with the help of (16) and (17) equation (9) reduces to

$$
\begin{equation*}
J\left(\mathbf{Q}_{0}, \omega_{0}\right)=A S\left(\mathbf{Q}_{0}, \omega_{0}\right) \cdot V_{I} \cdot V_{F} . \tag{23}
\end{equation*}
$$

Thus fitting a theoretical $S\left(\mathbf{Q}_{0}, \omega_{0}\right)$ to $J\left(\mathbf{Q}_{0}, \omega_{0}\right)$ involves only multiplying $S$ by $V_{I} . V_{F}$ and it is not necessary to unfold the data in this case. This method has been used to analyze results on quartz (Bauer, Jagodzinski, Dorner \& Grimm, 1971).

As an illustrative example, we have applied equation (23) to inelastic neutron scattering data from $\mathrm{Tb}_{2}\left(\mathrm{MoO}_{4}\right)_{3}$ as described by Axe, Dorner \& Shirane (1971). This is shown in Fig. 3. The scattering law, which describes a heavily damped phonon of a soft mode, is, to a good approximation, linear within the resolution volume. Moreover, the temperature is high enough compared to the energy transfers so that the scattering law can be taken to be symmetric for energy gain and loss. Therefore, the asymmetry in Fig. 3 is entirely due to the normalization of the resolution function. A least-squares fit using equation (23) gives excellent agreement with the experimental data.
(b) $S(\mathbf{Q}, \omega)$ planar in four-dimensional space

By planar we mean that (i) $S(\mathbf{Q}, \omega)$ is restricted in the dimension perpendicular to a plane, (ii) $S(\mathbf{Q}, \omega)$ is not curved within the resolution volume. For harmonic phonons with infinite lifetime the plane, which represents the dispersion surface, has effectively zero thickness, and this can be expressed by means of a $\delta$-function. When scanning through such an $S(\mathbf{Q}, \omega)$ one has to consider the variation of two quantities, the efficiency factor $R_{0}$ and the shape of the resolution function. The shape is given by the widths $w_{k}$. We assume that the orientations of the principal axes of $R$ do not change during the scan.

We must now consider how much of the change in the normalization $V_{I}, V_{F}$ along the scan is due to a change in the efficiency factor and how much is due to changes in the widths. To do this, we must refer back to equation (11), where we showed that at least as far as the $\mathbf{Q}$ dependence is concerned, $R$ is a convolution of $p_{i}\left(k_{i}\right)$ with $p_{f}\left(k_{f}\right)$. Simplifying, we describe both $p_{i}$ and $p_{f}$ by one-dimensional Gaussians with widths $\eta_{1}$ and $\eta_{2}$ respectively. The folding of these two Gaussians $G(x)$ now represents the resolution function $R$. In this expression $x$ stands for the four dimensions of $R$. Later, $x$ will represent enly the dimension perpendicular to the plare representing $S$.

$$
\begin{gather*}
G\left(x-x_{0}\right) \sim \frac{\eta_{1} \cdot \eta_{2}}{l\left(\eta_{1}^{2}+\eta_{2}^{2}\right)} \cdot \exp \left[-\frac{\left(x-x_{0}\right)^{2}}{2\left(\eta_{1}^{2}+\eta_{2}^{2}\right)}\right] .  \tag{24}\\
\int G\left(x-x_{0}\right) \mathrm{d} x \sim \eta_{1} \cdot \eta_{2} \equiv \text { normalization }
\end{gather*}
$$

$$
\frac{\eta_{1} \cdot \eta_{2}}{V\left(\eta_{1}^{2}+\overline{\eta_{2}^{2}}\right)} \equiv \text { efficiency factor } \sim \frac{\text { normalization }}{\text { width }}
$$

Let us assume $\eta_{1}^{2} \gg \eta_{2}^{2}$ or in other words that $p_{i}$ is much wider than $p_{f}$. Then the width of $G(x)$ is entirely determined by $\eta_{1}$ and the efficiency factor is $\eta_{2}$

$$
\begin{gather*}
G\left(x-x_{0}\right) \sim \eta_{2} \cdot \exp \left\{-\frac{\left(x-x_{0}\right)^{2}}{2 \eta_{1}^{2}}\right\} . \text {. } \quad \eta_{1}^{2} \gg \eta_{2}^{2} \tag{25}
\end{gather*}
$$

An inelastic scan can be made either by varying $k_{f}$ and holding $k_{I}$ fixed or vice versa. The first possibility, i.e. $k_{I}$ fixed, implies that $\eta_{1}$ is fixed. Hence, the width of the resolution is unchanged, during the scan and the 'overlapping' of $S$ and $R$ is the same on both sides of the maximum. Only the efficiency factor $\eta_{2}$ varies. Dividing the data at each point by $\eta_{2}$ (which in effect is dividing by $V_{F}$ ) is equivalent to a representation in which the scattering law $S$ is folded with a normalized resolution function (see section 6).

A second possibility is that $\eta_{2}$ be fixed while $\eta_{1}$ varies. This is more complicated because now the efficiency factor is constant and all of the variation is due to the change in the width of the resolution function. In this case we assume that $x$ represents only the dimension perpendicular to the plane. This means that the resolution function is already integrated over all three dimensions in the 'plane', or in other words that the resolution function is projected (see Section 6) onto an axis perpendicular to the plane.

We are interested in the value of $G$ at $x$, where $x$ is assumed to be the intersection with the plane, while the instrument is at a nominal position $x_{0}$. The width $\eta_{1}$ of $G$ depends on $x_{0}$. Expanding equation (25) in a Taylor series [with respect to $\eta_{1}$ around $\eta_{10}$ (at $\left.x=x_{0}\right)$ ] gives

$$
\begin{align*}
& G\left(x-x_{0}, \eta_{1}\right)=G\left(x-x_{0}, \eta_{10}\right)+\eta_{2} \\
& \times \exp \left\{-\frac{\left(x-x_{0}\right)^{2}}{2 \eta_{10}^{2}}\right\} \frac{\left(x-x_{0}\right)^{2}}{\eta_{10}^{3}}\left(\eta_{1}-\eta_{10}\right)+\ldots \tag{26}
\end{align*}
$$

For small $\left(\eta_{1}-\eta_{10}\right)$ we now obtain $G\left(x-x_{0}, \eta_{1}\right)$ for a scan which changes the nominal position $x_{0}$. The variation of $\eta_{1}$ with $x_{0}$ is not of particular interest in this context. Fig. 4 shows $G\left(x-x_{0}, \eta_{1}\right) / G\left(x-x_{0}, \eta_{10}\right)$ as a function of $\eta_{1}$. At the position $x-x_{0}=\eta_{10}$

$$
\begin{align*}
& G\left(\eta_{10}, \eta_{1}\right)  \tag{27}\\
& G\left(\eta_{10}, \eta_{10}\right)
\end{aligned}=1+\begin{aligned}
& \eta_{1} \\
& \eta_{10}
\end{align*}-1=\frac{\eta_{1}}{\eta_{10}} .
$$

In Fig. 4 the result of applying equation (27) appears as the broken line. It can be seen that equation (27) agrees with equation (26) at $x-x_{0}=0$, and $x-x_{0}=$ $\pm \eta_{10}$, i.e. at positions where $G\left(x, \eta_{10}\right)$ is 1 and $0 \cdot 61$. The broken line is the approximation to the correction obtained by simply dividing by $\eta_{1}$, which is equivalent to dividing by $V_{1}$.

Fig. 4 gives the impression that the disagreement between the true correction (the solid line) and the approximate correction (the broken line) is much larger than in fact it is for the four-dimensional resolution function. This is because we have been considering only one dimension. The integration over three dimensions within the plane is already done for every point by the instrument and the approximation there-
fore affects only one of the four dimensions of the resolution function.

The approximation is even better if $\eta_{1} \simeq \eta_{2}$ (i.e. $V_{I} \simeq V_{F}$ ), because the change in the width is smaller. The variation of $\eta_{1}$ will now affect both the width and the efficiency factor, and the change in the efficiency factor will be exactly taken into account by $V_{I}, V_{F}$ as was shown above.

Thus, for well defined phonons the method of dividing every measured point by $V_{I} . V_{F}$ is a good way to obtain data which represent a convolution of the phonon scattering law with a normalized resolution function (see Section 6).

A monitor with a $\frac{1}{k_{I}}$, characteristic, set in the incident monochromatic beam to control the counting time, directly corrects for the variation of $V_{1}$. Hence, for cases $5 a$ and $5 b$ with $k_{F}$ fixed and a $\frac{1}{k_{1}}$ monitor, no corrections for changing resolution are necessary.

## 6. Discussion of the data corrected by the normalization

In Section 5 we discussed how the normalization of the resolution function affects the intensity of each measured point. Now we will consider the measured widths and integrated intensities in more detail.

To do this, we will first show an interesting property of Gaussian functions for a two-dimensional example. This is most easily understood by imagining that $x$ is a coordinate perpendicular to the plane representing $S(\mathbf{Q}, \omega)$ and $y$ stands for the three coordinates in the


Fig. 5. The ellipses represent projections of four-dimensional ellipsoids for different probability. For a const- $Q_{0}$ scan the $Q_{t i}, \Theta_{t i}$ of contact with the scattering law $S(Q, \omega)$ are shown for four different $\omega_{0 i}$. Note that the lines connecting $Q_{0} \omega_{0}$ with $Q_{i t}, \omega_{t i}$ are parallel.
plane, as was done in Section 5. We mentioned there that the instrument performs an integration within the plane, i.e. along $y$. If we assume $S(\mathbf{Q}, \omega)$ to be constant, the integration can be performed over the resolution function alone for every $x$.

$$
\begin{align*}
& \int \exp \left\{-\left(a_{11} x^{2}+2 a_{12} x y+a_{22} y^{2}\right)\right\} \mathrm{d} y \\
& =\text { const. } \exp \left\{-\frac{a_{11} a_{22}-a_{12}^{2}}{a_{22}} \cdot x^{2}\right\} \tag{28}
\end{align*}
$$

At $x^{2}=\frac{a_{22}}{a_{11} a_{22}-a_{12}^{2}}$ the probability has dropped to $e^{-1}$ of the probability at $x=0$.

Now we compare the result of the above integration with the result of projecting the ellipse

$$
\begin{equation*}
a_{11} x^{2}+2 a_{12} x y+a_{22} y^{2}=1 \tag{29}
\end{equation*}
$$

onto the $x$ axis. We set $\frac{\mathrm{d} x}{\mathrm{~d} y}=0$ and calculate the coordinate $x_{p}$ and find

$$
\begin{equation*}
x_{p}^{2}=\stackrel{a_{22}}{a_{11} a_{22}-a_{12}} . \tag{30}
\end{equation*}
$$

Thus we see that a projection gives the same width as an integration. This holds for more than two dimensions as well.

Using the projection of the ellipse helps to visualize the factors affecting the measured intensity and line shape. In a scan, in which $\mathbf{Q}_{0}$ and $\omega_{0}$ vary, we define ellipsoids of different probabilities, so that for each value of $\mathbf{Q}_{0}, \omega_{0}$ one ellipsoid just contacts the plane representing $S(\mathbf{Q}, \omega)$ at $\mathbf{Q}_{t}, \omega_{t}$ (see Fig. 5). A scan perpendicular to the plane yields the smallest measured width.

Under the assumption that $S(\mathbf{Q}, \omega)$ is constant, the probability associated with the 'ellipsoid of contact' times $S$ gives the measured intensity also, the measured width for any direction of scan is larger than the smallest width by a factor $\frac{1}{\cos }-$ ( $\varphi$ being the angle between the scan and the normal to the plane). As long as $S$ is constant, a 'perpendicular' scan will not reveal more physical information than an inclined one.

A more realistic $S(\mathbf{Q}, \omega)$ varies with $\mathbf{Q}$ and $\omega$. Then to a good approximation, $S\left(\mathbf{Q}_{t}, \omega_{t}\right)$ at the contact point can be considered as giving the 'mean value' of $S(\mathbf{Q}, \omega)$, which contributes to the scattering at a nominal setting, $\mathbf{Q}_{0}, \omega_{0}$ (see Fig. 5). Therefore, the counting rate at $\mathbf{Q}_{0}, \omega_{0}$ is $S\left(\mathbf{Q}_{t}, \omega_{t}\right)$ times the probability on the surface of the contacting ellipsoid. For most scans $\mathbf{Q}_{t}, \omega_{t}$ will change with the variation of $\mathbf{Q}_{0}, \omega_{0}$. At the position of maximum overlap between $\mathbf{S}$ and $\mathbf{R}$ it follows that $\mathbf{Q}_{\mathrm{t}}=\mathbf{Q}_{0}$ and $\omega_{t}=\omega_{0}$.

We understand that the 'mean value' of $S(\mathbf{Q}, \omega)$ described above will change during a scan varying $\mathbf{Q}_{0}, \omega_{0}$. If $S(\mathbf{Q}, \omega)$ is known approximately, one can
evaluate the values of $\mathbf{Q}_{t}, \omega_{t}$ and take into account the varying 'mean value' of $S(\mathbf{Q}, \omega)$. This correction has been performed for phonon measurements in argon (Dorner \& Egger, 1971) and in CD $_{4}$ (Press, Dorner \& Stiller, 1971). For $S(\mathbf{Q}, \omega)$ described by a plane, all $\mathbf{Q}_{t}, \omega_{t}$ along the scan lie on a straight line through the center of $R$. Having taken into account both the 'mean value' of $S$ and the normalization corrections explained in Section 5, the unfolding reduces to a one-dimensional problem even if the phonon has an intrinsic line shape, since the contribution of the resolution to the measured width is known. A favourable scan is a scan which has the same $\mathbf{Q}_{t}, \omega_{t}$ for all $\mathbf{Q}_{0}, \omega_{0}$.

Consider now the measured intensities $I\left(\mathbf{Q}_{0}, \omega_{0}\right)$ corrected by $V_{I}\left[k_{l}\left(\mathbf{Q}_{0}, \omega_{0}\right)\right]$ and $V_{F}\left[k_{F}\left(\mathbf{Q}_{0}, \omega_{0}\right)\right]$. We rewrite equation (9)

$$
\begin{align*}
& \frac{I\left(\mathbf{Q}_{0}, \omega_{0}\right)}{V_{I}, V_{F}}=\frac{J\left(\mathbf{Q}_{0}, \omega_{0}\right)}{V_{I} \cdot \bar{V}_{F}} \cdot \Delta t=A \cdot \Delta t \int S(\mathbf{Q}, \omega) \\
& \times \frac{R\left(\mathbf{Q}-\mathbf{Q}_{0},\left(\omega-\omega_{0}\right)\right.}{V_{I} \cdot V_{F}} \mathrm{~d}^{3} Q \mathrm{~d} \omega \tag{31}
\end{align*}
$$

where $\Delta t$ is the measuring time which is assumed to be the same for each point. Since $R\left(\mathbf{Q}-\mathbf{Q}_{0}, \omega-\omega_{0}\right)$ depends on the zero points $\mathbf{Q}_{0}$ and $\omega_{0}$, an integration of equation (9) over $\mathbf{Q}_{0}$ and $\omega_{0}$ would not be easy. But, $R\left(\mathbf{Q}-\mathbf{Q}_{0}, \omega-\omega_{0}\right)$
$V_{I}, V_{F}$ is a normalized function independent of $\mathbf{Q}_{0}, \omega_{0}$ within the approximation discussed in $5 b$. Since this integral is independent of $\mathbf{Q}$ and $\omega$, we integrate equation (31) over $\mathrm{d} \mathbf{Q}_{0} \cdot \mathrm{~d} \omega_{0}$ and get

$$
\begin{equation*}
\int \frac{I\left(\mathbf{Q}_{0}, \omega_{0}\right)}{V_{1} \cdot V_{F}} \mathrm{~d} \mathbf{Q}_{0} \mathrm{~d} \omega_{0}=A \cdot \Delta t \cdot \int \mathrm{~S}(\mathbf{Q}, \omega) \mathrm{d}^{3} Q \mathrm{~d} \omega \tag{32}
\end{equation*}
$$

Equation (32) shows us that the integration (effectively a summation) over the corrected intensities yields the integral over the scattering law.

Hence $\frac{I\left(\mathbf{Q}_{0}, \omega_{0}\right)}{V_{1}, V_{F}}$ is a folding of $S(\mathbf{Q}, \omega)$ with a resolution function normalized to unity. The unfolding of a Gaussian with $S(\mathbf{Q}, \omega)$ will not be further discussed here.

## 7. Conclusion

We have derived an exact normalization of the resolution function and have shown that the so-called efficiency factor $R_{0}$ can be expressed as a function of the normalization and the widths in the direction of the principal axes. This simple result for $R_{0}$ should be a great help for numerical evaluation of the folding of the resolution function with the scattering law.

Assuming that $S(\mathbf{Q}, \omega)$ varies linearly over the volume of $R$ (at least with respect to three dimensions) we give a method for correcting the measured data without numerical convolution. In the case of phonons it is necessary to assume that $S(\mathbf{Q}, \omega)$ is planar to apply this technique. In this case, we show that the unfolding of the data reduces to a one-dimensional problem and
the integral over the corrected data yields the integral over the scattering law.

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## APPENDIX

Starting with Cooper \& Nathans's (1967) equation (6) we find

$$
\begin{align*}
& p_{i}\left(\mathbf{k}_{i}\right) \cdot \mathrm{d} k_{i x} \cdot \mathrm{~d} k_{i y} \cdot \mathrm{~d} k_{i z}=P_{M}\left(k_{I}\right) P_{0 M} \\
& \quad \times \exp \left\{\frac { 1 } { 2 } \left[\left(\frac{\left(\Delta k_{i} / k_{I}\right) \tan \theta_{M}+\gamma_{1}}{\eta_{M}}\right)^{2}\right.\right. \\
& \quad+\left(\frac{2\left(\Delta k_{i} / k_{I}\right) \tan \theta_{M}+\gamma_{1}}{\alpha_{0}}\right)^{2}+\frac{\gamma_{1}^{2}}{\alpha_{1}^{2}} \\
& \left.\left.\quad+\left(4 \sin ^{2}-\frac{1}{\theta_{M} \eta_{M}^{2}+\beta_{0}^{2}}+\frac{1}{\beta_{1}^{2}}\right) \delta^{2}\right]\right\} \\
& \quad \times \mathrm{d}\left(\Delta k_{i}\right) \cdot k_{I} \cdot \mathrm{~d} \gamma_{1} \cdot k_{I} \mathrm{~d} \delta_{1} \tag{33}
\end{align*}
$$

where $\mathrm{d}\left(\Delta k_{i}\right)$ is the component in the direction of $\mathbf{k}_{1}, k_{1} \cdot \mathrm{~d} \gamma_{1}$ is the component in the scattering plane perpendicular to $k_{I}$ and $k_{I} \cdot \mathrm{~d} \delta_{I}$ is the component perpendicular to the scattering plane.
Equation (6) of Cooper \& Nathans contains several minor errors. First, the denominator of the $\delta_{1}$ and $\delta_{2}$ expressions should contain $\sin \theta_{M}$ and $\sin \theta_{A}$ instead of $\tan \theta_{M}$ and $\tan \theta_{A}$. Second, they include only one $P_{o}$, which is derived in their Appendix I for the monochromator; a similar expression for the analyzer is missing.

There is a more serious objection to their expression for $P_{o}$, as given in their equation (24). If we assume $\beta_{1} \gg \beta_{0}$ and $\eta_{M}^{\prime}$, then all incoming neutrons with vertical divergence defined by $\beta_{0}$ must be reflected by the monochromator. To understand how this comes about, we note that the attenuation produced by the finite reflectivity of the monochromator is expressed by $P_{M}$ [actually $P_{M}\left(k_{I}\right)$ ]. The vertical collimation plays no role in determining the values of $k_{i}$, which will be reflected; this is entirely determined by the horizontal Bragg condition. Thus, if $P_{M}\left(k_{I}\right)$ is set equal to unity, we should expect that the integral over $\delta_{1}$ is proportional to $\beta_{0}$ or in other words that every neutron with an acceptable $k_{i}$ and a vertical divergence within $\beta_{0}$ will be reflected. In fact, this condition is not satisfied if Cooper \& Nathans's formula (24) is integrated. This consistency consideration has been used to derive correct expressions for $P_{0 M}$ and $P_{0 A}$, which are

$$
\begin{align*}
& P_{0 M}=\frac{\beta_{0}}{l^{\prime} 4 \sin ^{2}} \frac{\left(\theta_{M}\right) \eta_{M}^{2}+\beta_{0}^{2}}{}, \\
& P_{0 A}=\frac{\beta_{3}}{14 \sin ^{2}\left(\theta_{A}\right) \eta_{A}^{\prime 2}+\beta_{3}^{2}} \tag{34}
\end{align*}
$$

Chesser (1971) has pointed out that this error is a consequence of the fact that Cooper and Nathans's equation (14), which describes the probability of vertical broadening of the beam by reflexion from a crystal with a vertical mosaic spread, is not normalized. Equation (C \& N 14) should read
$P\left(\delta_{1}-\delta_{0}\right)=V^{\prime}(2 \pi) \cdot 2 \frac{1}{\sin } \bar{\theta}_{M} \eta_{M}^{\prime} \exp \left\{-\frac{1}{2} k_{I}^{2}\binom{\delta_{1}-\delta_{0}}{Q_{M} \cdot \eta_{M}^{\prime}}^{2}\right\}$.
(C \& N 14)
We used the Bragg equation 2. $k_{I} \cdot \sin \theta_{M}=Q_{M}$.
The correct $P_{0 M}$ consequently reads

$$
\begin{equation*}
P_{0 M}=\frac{1}{l^{\prime}(2 \bar{\pi}) \cdot 2 \sin \theta_{M} \eta_{M}^{\prime}} \cdot P_{0 C \& N} \tag{35}
\end{equation*}
$$

where $P_{0 C \& N}$ is given by Cooper and Nathans's equation (24).

Comparing our $P_{0 M}$ with equation (5.3) of Tucciarone et al. (1971) we found a misprint in their equation (5.3), as a factor $2 \pi \cdot \eta_{M}^{\prime} \cdot \eta_{A}^{\prime}$ is missing in the denominator. Integrating equation (33) we get

$$
\begin{aligned}
& V_{I}=P_{M}\left(k_{I}\right) \cdot(2 \pi)^{3 / 2} \times \\
& \cdots \\
& \tan \theta_{M} \cdot\left[\left(\begin{array}{c}
1 \\
\eta_{M}^{2}
\end{array}+\frac{4}{\alpha_{0}^{2}}\right)\left(\begin{array}{c}
1 \\
\eta_{M}^{2}
\end{array}+\frac{1}{\alpha_{0}^{2}}+\frac{1}{\alpha_{1}^{2}}\right)-\left(\begin{array}{c}
1 \\
\eta_{M}^{2}
\end{array}+\frac{2}{\alpha_{0}^{2}}\right)^{2}\right]^{1 / 2} \\
& \times \sqrt{\left(4 \sin ^{2}\left(\theta_{M}\right) \overline{\eta_{M}^{\prime 2}}+\beta_{0}^{2}\right)} \\
& \times \sqrt{\left(4 \sin ^{2}\left(\theta_{M}\right) \eta_{M}^{\prime 2}+\beta_{0}^{2}+\frac{1}{\beta_{1}^{2}}\right)} \cdot k_{I}^{2}
\end{aligned}
$$

This reduces to

$$
\begin{align*}
V_{I}= & P_{M}\left(k_{I}\right) \cdot(2 \pi)^{3 / 2} \cdot k_{I}^{3} \cot \theta_{M} \\
& \times \frac{\beta_{0} \cdot \beta_{1}}{\left.\sqrt{\left(4 \sin ^{2}\right.}\left(\theta_{M}\right) \eta_{M}^{\prime 2}+\beta_{0}^{2}+\beta_{1}^{2}\right) \cdot} \cdot \sqrt[\eta_{M} \alpha_{0} \alpha_{1}]{\left.\sqrt{\left(\alpha_{0}^{2}+\alpha_{1}^{2}+4 \eta_{M}^{2}\right.}\right)} . \tag{37}
\end{align*}
$$

A similar expression can also be derived for the analyzer, where $P_{A}\left(k_{F}\right)$ represents the counter efficiency and the reflectivity of the analyzing crystal.

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[^0]:    * Work performed under the auspices of the U.S. Atomic Energy Commission.
    $\dagger$ Permanent address.
    $\ddagger$ By Gaussian approximation, we mean that Gaussian distributions are used for all quantities which determine the properties of the neutron beams, such as collimators, mosaic widths, pulse lengths, energy widths, counter thicknesses, etc.

[^1]:    § In the course of deriving the normalization we discovered an error in formula (24) of Cooper \& Nathans. This is discussed in Appendix I.

